

Critical magnetic fields for the magnetic Dirac-Coulomb operator

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Earth's magnetic field = 1 Gauss

Maximal field on Earth (MRI) = 1 Tesla = 10^4 Gauss

Deatly field strength = 10^5 Tesla = 10^9 Gauss

In new magnetars one expectes fields of 10^{11} Tesla. In recent theories, up to 10^{16} Tesla in the heart of the magnetars.

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- hydrogenic atoms in a curved universe (Nowotny),
- intense electrostatic fields.... (... , Brodsky, Mohr,...),
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REMARK. Same problem but very different situation in nonrelativistic quantum mechanics (with the Schrödinger and the Pauli operators).

The free Dirac operator

$$\mathbf{H} = -\mathbf{i}\boldsymbol{\alpha} \cdot \nabla + \beta, \quad \alpha_1, \alpha_2, \alpha_3, \beta \in \mathcal{M}_{4 \times 4}(\mathbb{C}) \quad (c = m = \hbar = 1)$$

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QUESTION : Spectrum of $\mathbf{H} + V$?

Eigenvalues of the Dirac operator without magnetic field

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Its spectrum is given by:

$$\sigma(H_\nu) = (-\infty, -1] \cup \{\lambda_1^\nu, \lambda_2^\nu, \dots\} \cup [1, \infty)$$

$$0 < \lambda_1^\nu = \sqrt{1 - \nu^2} \leq \dots \leq \lambda_k^\nu \dots < 1.$$

and the fact that $\lambda_1(H + V_\nu)$ belongs to $(-1, 1)$ is a kind of "stability condition" for the electron.

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Does $\lambda_1(B, V)$ ever leave the spectral gap $(-1, 1)$? and if yes, for which values of B ?

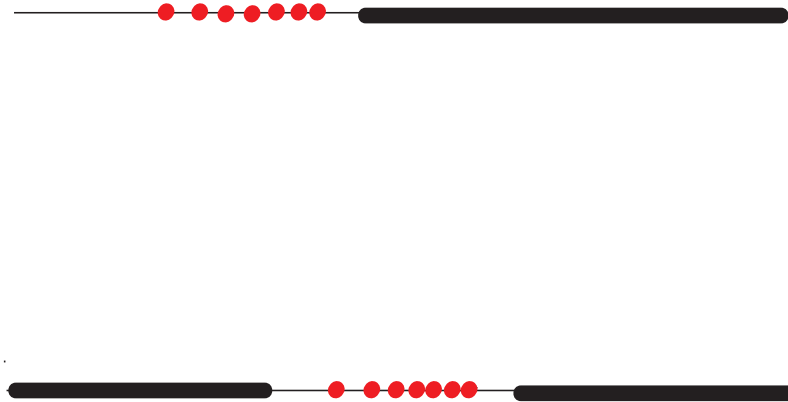
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In the second case, $\lambda_1 = \min \max \frac{(Ax, x)}{\|x\|^2}$, $\lambda_1 = \max \min \frac{(Ax, x)}{\|x\|^2}$, \dots

Abstract min-max theorem (Dolbeault, E., Séré, 2000)

Let \mathcal{H} be a Hilbert space and $A : F = D(A) \subset \mathcal{H} \rightarrow \mathcal{H}$ a self-adjoint operator defined on \mathcal{H} . Let \mathcal{H}_+ , \mathcal{H}_- be two orthogonal subspaces of \mathcal{H} satisfying: $\mathcal{H} = \mathcal{H}_+ \oplus \mathcal{H}_-$. Define $F_{\pm} := P_{\pm}F$.

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If (ii) $c_1 > a_-$, then c_k is the k -th eigenvalue of A in the interval (a_-, b) , where $b = \inf(\sigma_{\text{ess}}(A) \cap (a_-, +\infty))$.

Application to magnetic Dirac operators I

$$\psi : \mathbb{R}^3 \rightarrow \mathbb{C}^4, \quad \psi = \begin{pmatrix} \varphi \\ \chi \end{pmatrix} = \begin{pmatrix} \varphi \\ 0 \end{pmatrix} + \begin{pmatrix} 0 \\ \chi \end{pmatrix}, \quad \varphi, \chi : \mathbb{R}^3 \rightarrow \mathbb{C}^2$$

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Then, for all $k \geq 1$,

$$\lambda_k(B, V) = \inf_{\substack{Y \text{ subspace of } C_0^\infty(\mathbb{R}^3, \mathbb{C}^2) \\ \dim Y = k}} \sup_{\varphi \in Y \setminus \{0\}} \sup_{\substack{\psi = \begin{pmatrix} \varphi \\ \chi \end{pmatrix} \\ \chi \in C_0^\infty(\mathbb{R}^3, \mathbb{C}^2)}} \frac{(\psi, (H_B + V)\psi)}{(\psi, \psi)}$$

Application to magnetic Dirac operators II

The first eigenvalue of $H_B + V$ in the spectral hole $(-1, 1)$ is given by

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is the unique real number λ such that

$$\int_{\mathbb{R}^3} \left(\frac{|\sigma \cdot \nabla_B \varphi|^2}{1 - V + \lambda} + (1 + V)|\varphi|^2 \right) dx = \lambda \int_{\mathbb{R}^3} |\varphi|^2 dx .$$

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$$\int_{\mathbb{R}^3} \frac{|(\sigma \cdot \nabla_B) \varphi|^2}{1 + \lambda_1(B, V_\nu) + \frac{\nu}{|x|}} dx + (1 - \lambda_1(B, V_\nu)) \int_{\mathbb{R}^3} |\varphi|^2 dx \geq \int_{\mathbb{R}^3} \frac{\nu}{|x|} |\varphi|^2 dx$$

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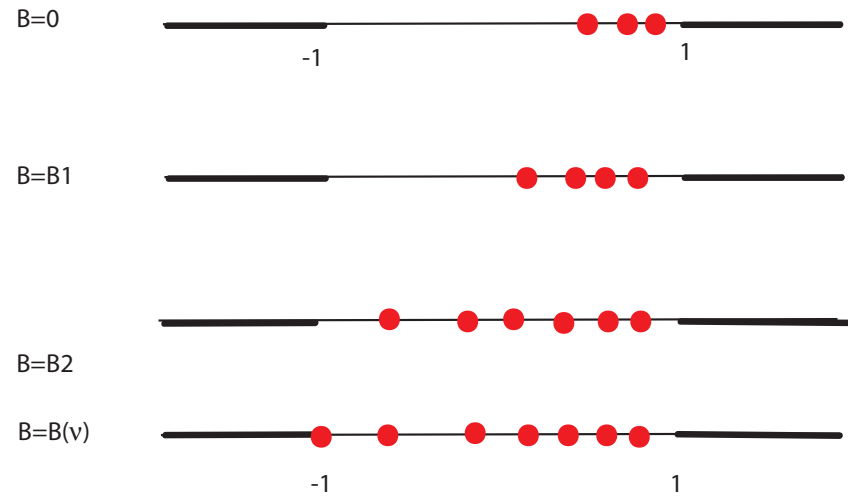
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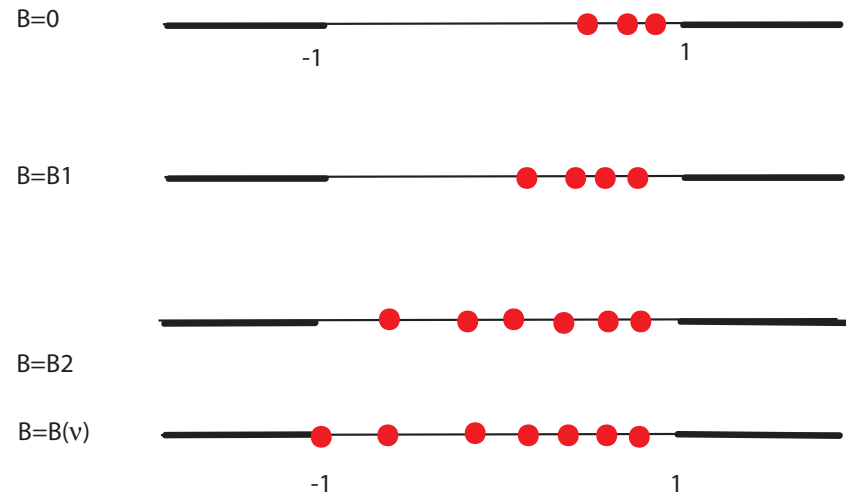
QUESTIONS : When do we have $\lambda_1(B, V) \in (-1, 1)$?

If the eigenvalue $\lambda_1(B, V)$ leaves the interval $(-1, 1)$, when ?

For a potential $V = V_\nu$, $\nu \in (0, 1)$, $0 < B_1 < B_2 < B(\nu)$:

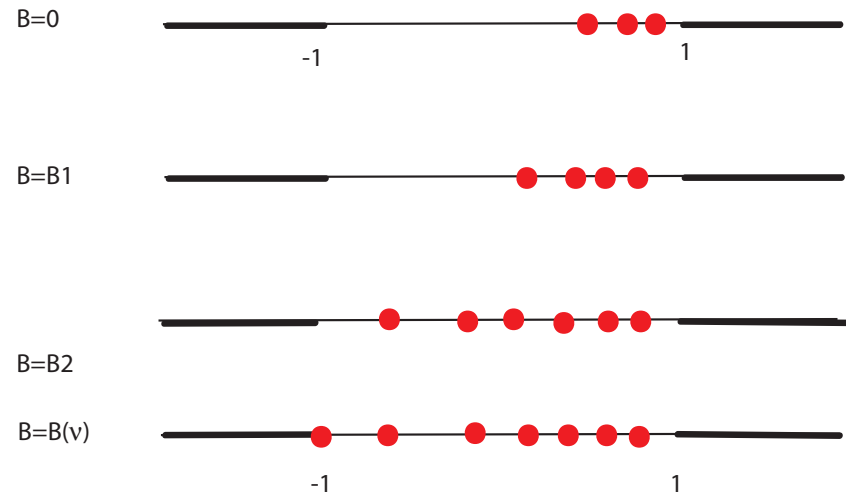


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DEFINITION: $B(\nu) := \inf \{ B > 0 : \liminf_{b \nearrow B} \lambda_1(B, \nu) = -1 \} .$

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TOOL TO ESTIMATE IT :

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First results to estimate $\lambda_1(B, V)$

For a given \bar{C} , we have $\lambda_1(B, V) < \bar{C}$ if we can find a function $\bar{\varphi}$ s. t.

$$(*) \quad \int_{\mathbb{R}^3} \frac{|(\sigma \cdot \nabla_B) \bar{\varphi}|^2}{1 + \bar{C} - V} dx + (1 - \bar{C}) \int_{\mathbb{R}^3} |\bar{\varphi}|^2 dx + \int_{\mathbb{R}^3} V |\bar{\varphi}|^2 dx < 0$$

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And to verify that $\lambda_1(B, V) > \underline{C}$ it is enough to prove that for all φ

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and of course we are interested in the case $\overline{C} = 1$ in (*), and $\underline{C} = -1$ in (**).

Some results (with J. Dolbeault and M. Loss)

$$V_\nu = -\frac{\nu}{|x|}, \quad \mathbf{A}_B(x) := \frac{B}{2} \begin{pmatrix} -x_2 \\ x_1 \\ 0 \end{pmatrix}, \quad \mathbf{B}(x) := \begin{pmatrix} 0 \\ 0 \\ B \end{pmatrix}; \quad \nu \in (0, 1), \quad B \geq 0$$

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- $\lim_{\nu \rightarrow 1} B(\nu) > 0$, $\lim_{\nu \rightarrow 0} \nu \log B(\nu) = \pi$

Some results (with J. Dolbeault and M. Loss)

$$V_\nu = -\frac{\nu}{|x|}, \quad \mathbf{A}_B(x) := \frac{B}{2} \begin{pmatrix} -x_2 \\ x_1 \\ 0 \end{pmatrix}, \quad \mathbf{B}(x) := \begin{pmatrix} 0 \\ 0 \\ B \end{pmatrix}; \quad \nu \in (0, 1), \quad B \geq 0$$

- For $B = 0$, $\lambda_1(0, V_\nu) = \sqrt{1 - \nu^2} \in (-1, 1)$
- For all $B \geq 0$, $\lambda_1(B, V_\nu) < 1$
- For all $\nu \in (0, 1)$ there exists a critical field strength $B(\nu)$ such that

$$\lambda_1(B, V_\nu) \leq -1 \quad \text{if} \quad B \geq B(\nu)$$

- $\lim_{\nu \rightarrow 1} B(\nu) > 0$, $\lim_{\nu \rightarrow 0} \nu \log B(\nu) = \pi$
- For ν small, the asymptotics of $B(\nu)$ can be calculated by an approximation in the first relativistic “Landau level”.

A better method to determine $B(\nu)$

$$\int_{\mathbb{R}^3} \frac{|(\sigma \cdot \nabla_B)\varphi|^2}{1 + \lambda_1(B, V_\nu) + \frac{\nu}{|x|}} dx + (1 - \lambda_1(B, V_\nu)) \int_{\mathbb{R}^3} |\varphi|^2 dx \geq \int_{\mathbb{R}^3} \frac{\nu}{|x|} |\varphi|^2 dx$$

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and we are looking for $B_n \rightarrow B(\nu)$ such that $\lambda_1(B_n, \nu) \rightarrow -1$.

If everything were compact, we would be able to pass to the limit and obtain

$$\int_{\mathbb{R}^3} \frac{|x| |(\sigma \cdot \nabla_B)\varphi|^2}{\nu} - \int_{\mathbb{R}^3} \frac{\nu}{|x|} |\varphi|^2 dx + 2 \int_{\mathbb{R}^3} |\varphi|^2 dx \geq 0$$

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Now define the functional

$$\mathcal{E}_{B,\nu}[\phi] := \int_{\mathbb{R}^3} \frac{|x|}{\nu} |(\sigma \cdot \nabla_B)\phi|^2 dx - \int_{\mathbb{R}^3} \frac{\nu}{|x|} |\phi|^2 dx ,$$

If everything were compact and "nice",

$$\mu_{B,\nu} + 2 = 0 ; \quad \mu_{B,\nu} := \inf \left\{ \mathcal{E}_{B,\nu}[\phi] ; \int_{\mathbb{R}^3} |\varphi|^2 dx = 1 \right\} .$$

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Now we would like to estimate $B(\nu)$. This can be done analytically or/and numerically.

As we said before, analytically we have some estimates for ν close to 0 and to 1.

The Landau level approximation

Consider the class of functions $\mathcal{A}(B, \nu)$:

$$\phi_\ell := \frac{B}{\sqrt{2\pi} 2^\ell \ell!} (x_2 + i x_1)^\ell e^{-B s^2/4} \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \ell \in \mathbb{N}, \quad s^2 = x_1^2 + x_2^2,$$

where the coefficients depend only on x_3 , i.e.,

$$\phi(x) = \sum_{\ell} f_\ell(x_3) \phi_\ell(x_1, x_2), \quad z := x_3.$$

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Now, we shall restrict the functional $\mathcal{E}_{B, \nu}$ to the first Landau level. In this framework, that we shall call the **Landau level ansatz**, we also define a critical field by

$$B_{\mathcal{L}}(\nu) := \inf \left\{ B > 0 : \liminf_{b \nearrow B} \lambda_1^{\mathcal{L}}(b, \nu) = -1 \right\},$$

where

$$\lambda_1^{\mathcal{L}}(b, \nu) := \inf_{\phi \in \mathcal{A}(B, \nu), \Pi^\perp \phi = 0} \lambda[\phi, b, \nu].$$

THEOREM : $B^{\mathcal{L}}(\nu) = \frac{4}{\mu^{\mathcal{L}}(\nu)^2}$, where

$$\mu^{\mathcal{L}}(\nu) := \inf_f \frac{\mathcal{L}_\nu[f]}{\|f\|_{L^2(\mathbb{R}^+)}^2} < 0.$$

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COROLLARY. $\mu(\nu) \leq \mu^{\mathcal{L}}(\nu) < 0 \implies B(\nu) \leq B^{\mathcal{L}}(\nu).$

THEOREM. For $\nu \in (0, \nu_0)$, $B^{\mathcal{L}}(\nu + \nu^{3/2}) \leq B(\nu) \leq B^{\mathcal{L}}(\nu)$

Final results

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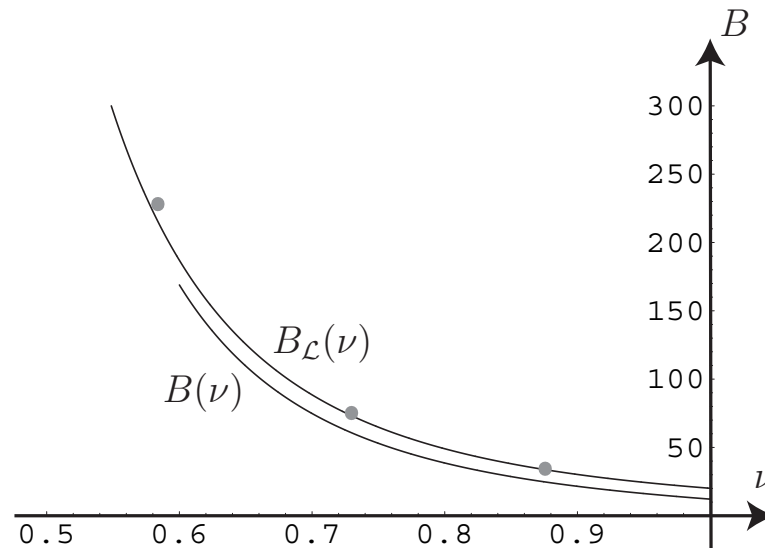
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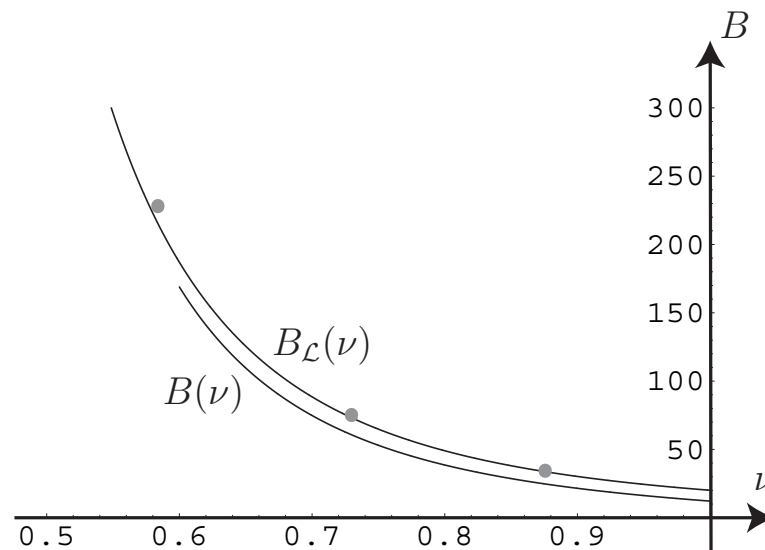


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NUMERICAL OBSERVATION. For ν near 1, $B(\nu)$ is below $B^{\mathcal{L}}(\nu)$ by 30%.

Dear John :

even if late ,

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**BON ANNIVERSAIRE !
HAPPY BIRTHDAY !**